

Experimental study of resonant sates in ²⁶Si and ²⁷P via elastic scattering of ²⁵Al+p and ²⁶Si+p

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The rapid proton capture process (rp process) is a dominant reaction sequence in explosive hydrogen burning that synthesizes heavier elements, especially proton-rich unstable nuclides. Therefore, accurate thermonuclear reaction rates for (p,γ) reactions on the rp process path are essential for an understanding of the nucleosynthesis processes and energy production. We studied proton resonant states in ²⁶Si and ²⁷P via elastic scattering in inverse kinematics at the low-energy RI beam facility CRIB (CNS Radioactive Ion Beam separator), University of Tokyo. By using a thick H₂ gas target, excitation energies range of 6.8 to 8.2 MeV in ²⁶Si and 2.3 to 3.8 MeV in ²⁷P were scanned, respectively. Several resonances above the proton threshold level were observed with high statistics and free from any background contribution in the target. Their resonance parameters were subsequently extracted by an R-matrix analysis; these are important to better constrain the production rates of the ²⁵Al(p,\gamma)²⁶Si and ²⁶Si(p,\gamma)²⁷P reactions.

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1. Introduction

The ²⁵Al(p, γ)²⁶Si and ²⁶Si(p, γ)²⁷P reactions are relevant to the synthesis of galactic ²⁶Al^{gs} (t_{1/2} = 0.717 million years), which decays to the first excited state in ²⁶Mg, giving rise to a 1.809-MeV γ -ray [1]. In principle, internal transitions from the isomer to the ground state of ²⁶Al are forbidden by their large nuclear spin difference as shown in figure 1. However, if the ²⁶Al^m is effectively communicated with the ²⁶Al^{gs} through the thermal population of higher excited levels above the critical temperature of T = 0.4 GK [2-4], then this thermal process whould enhance the production of ²⁶Al^{gs}. The ²⁵Al(p, γ)²⁶Si reaction becomes significantly faster than the β decay of ²⁶Al at higher temperatures [5]. Moreover the ²⁶Si(p, γ)²⁷P reaction competes with the β decay of ²⁶Al^m. Therefore the destruction of ²⁶Si is an important consideration, and reaction rates of ²⁵Al(p, γ)²⁶Si and ²⁶Si(p, γ)²⁷P should be determined accurately for an estimation of the production of ²⁶Al in explosive hydrogen burning.

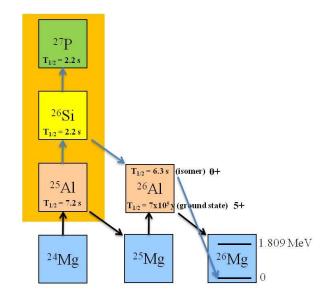


Figure 1: Reaction paths towars the production of the ²⁶Al

Apart from considerations related to ²⁶Al, these reactions may play a role in novae of the hotter variety or type I X-ray bursts. In such scenarios processing occurs mostly via the rp process which control the energy generation and nucleosystems. Model simulations show that these reactions are activated along the rp-process paths in explosive hydrogen burning [6,7].

The thermonuclear reaction rates of stellar capture reactions are determined by the properties of energy levels in the compound nucleus within the Gamow windows corresponding to the stellar temperatures at which these reactions occur. Therefore states in ²⁶Si and ²⁷P need to be well understood in order to determine these proton capture reaction rates. The uncertainty in these rates is mainly due to the lack of nuclear structure information above the proton threshold. States in ²⁶Si have been studied with different reactions, for example: (p,t), (³He,n), (³He,⁶He), (⁴He, ⁶He), and (p,p). Although many states have been discovered, their parameters are still

unmeasured at energies corresponding to higher temperatures. Moreover, a comparison with its mirror nucleus ²⁶Mg reveals that missing states still remain. Concerning the structure of ²⁷P, it has been studied with measurements of the (³He,⁸Li) and (⁷Li,⁸He) reactions, and Coulomb dissociation of ²⁷P. Even though five resonant states above the proton threshold were observed, the knowledge of the structure of ²⁷P is still insufficient because of uncertain resonance parameters such as resonance energies and spin-parity assignments. In addition, there is the possibility of finding a number of experimentally unobserved states.

2. Experiment

The measurments of ²⁵Al+p and ²⁶Si+p elastic scattering were performed at the low-energy RI beam facility CRIB [8,9] by impinging ²⁵Al and ²⁶Si radioactive ion beams onto a thick H₂ gas target in inverse kinematics. A ²⁴Mg primary beam with an energy of 7.5 MeV/nucleon bombarded a ³He gas target, where ²⁵Al and ²⁶Si beams were produced by the ³He(²⁴Mg,²⁵Al)d and ³He(²⁴Mg,²⁶Si)n reactions simultaneously, and separated by CRIB using the in-flight method. The secondary beams are contaminated with other isotopes including the ²⁴Mg primary beam. Beams were identified with a Time-Of-Flight (TOF) method and two Parallel Plate Avalanche Counters located upstream from the secondary target. The experimental details and results of ²⁶Si+p were published elsewhere [10]; the present article focuses on the results of ²⁵Al+p.

The radioactive ²⁵Al beam, with a well-defined energy of 71.58±0.54 MeV on target, an intensity of 3.3×10^4 pps, and a purity of 62%, was delivered to the secondary target position. The beam energy was measured with a silicon detector directly before the measurement of scattered protons. Protons elastically scattered from the ¹H(²⁵Al,p)²⁵Al reaction to forward angles in the laboratory frame were detected by using silicon detectors for a Δ E-E telescope. By calculating the elastic scattering kinematics of ²⁵Al+p, the measured proton energy of each event was converted to a center-of-mass energy, E_{c.m.}, corrected for the energy loss of particles in the target. The experimental differential cross sections of the proton scattering events were calculated from the number of selected proton events and incident beam ions, the target thickness, and the different solid angles depending on the interaction position in the thick target. Figure 2 is the excitation function at $\theta_{lab} = 0^{\circ}$ in E_{c.m.} Several peak-like structures were clearly observed. Corresponding excitation energies in ²⁶Si were calculated through the relation that E_x = E_{c.m.} + 5.518 (MeV) [11].

3. Data analysys and results

The differential cross section data has been analyzed using the R-matrix code (SAMMY-8.0.0) [12] to extract resonance parameters such as excitation energy E_x , spin *J*, parity π , and proton partial width Γ_p of resonance states. The spin-parity of the proton is $J^{\pi} = 1/2^+$ and that of the ground state of 25 Al is $5/2^+$, which couple together to give the incident channel spin S = 1/2 $\bigoplus 5/2 = 2$ or 3. Therefore, allowed spin-parities of the compound nucleus, 26 Si, are assigned as $J^{(-1)^1}$ where $J = l \bigoplus S$ and *l* is the relative orbital quantum number of the proton with respect to the nucleus.

Since at low energies below ~1.5 MeV Coulomb scattering is dominant or the widths of the resonances are expected to be too narrow, no clear resonance was observed in the excitation function over this energy region. The lowest resonance peak observed around 1.63 MeV was reproduced well by the R-matrix calculation with a $\mathcal{J}^{\pi} = 2^+$ assignment. This is in good agreement with previously reported assignment for this resonance level [13-15]. The second small peak around 1.89 MeV was fitted best with a $\mathcal{J}^{\pi} = 4^+$ assignment. The most recent study by Chen et al. [16] suggested $\mathcal{J}^{\pi} = 2^+$ assignment, however this is difficult to confirm due to the large error bars. As shown in figure 2(a), the result of fitting our experimental data with a J^{π} 2^+ assignment was worse than in case of a $J^{\pi} = 4^+$, the previously reported assignment [13]. A resonance to attempt to describe a small bump around 2.35 MeV in the middle of the two prominent peaks was included in the R-matrix calculation. Exclusion of this bump resulted in a less satisfactory fit for the tail of the most prominent peak around 2.14 MeV as shown in figure 2(a). An assignment of $J^{\pi} = 1^{-}$ for this state was adopted from previous work [14,15]. However we could not confirm the spin assignment clearly due to the narrow peak for this resonance. The third resonance (around 1.97 MeV), fits best with $J^{\pi} = 2^+$, which together with the level energy is in good agreement with the previously reported assignment for this level from transfer reaction studies [13-15] and the previous proton resonant scattering experiment [16].

As shown in figure 2(b), the fourth resonance at 2.14 MeV was fitted with $J^{\pi} = 2^+$, 3^+ or 3^- assignments, which were based on previous measurements [13-16]. The best fit was obtained for $J^{\pi} = 2^+$ or 3^+ , but $J^{\pi} = 3^-$ could not be reproduced and totally deviated from our data. Because the fit results of both $J^{\pi} = 2^+$ and 3^+ were consistent with experimental data, we could not determine the final spin assignment clearly for this resonance. Furthermore, in the recent high-resolution study of Matic et al. [13], two states in this region were measured at $E_x = 7.661(12)$ and 7.701(12) MeV, but there was no difference in fitting results between a single and a doublet state with our data.

Lastly, the fifth resonance around 2.46 MeV corresponds to a level at 2.501 (14) MeV from Chen et al. [16], which has observed this state firstly, suggesting an assignment of $J^{\pi} = 3^+$. However, there was almost no difference between $J^{\pi} = 2^+$ or 3^+ with our result, as shown in figure 2(c). Figure 2(d) shows best-fit results for the excitation function of ${}^{25}\text{Al}+p$ elastic scattering. The corresponding level energies and spin-parities in comparison with those from previous studies are given in Table 1. The uncertainty in the energy includes both systematic and fitting uncertainty.

this work	$^{1}\text{H}(^{25}\text{Al,p})$ [16]	(p,t) [13]
7.147(27), 2+	7.162(14), 2+	7.151(5)
7.401(28), 4+	7.402(40), 2+	7.415(23),(4+)
$7.484(28), 2^+$	7.484(13), 2+	7.479(12)
7.654(29), (2 ⁺ ,3 ⁺)		7.661(12),(2+)
	7.704(13), 3 ⁺	7.701(12)

Table 1: Level energies, E_x (MeV) and spin-parities in ²⁶Si in the range of 7.1 MeV to 8.3 MeV from the present work and in comparison with those of previous studies.

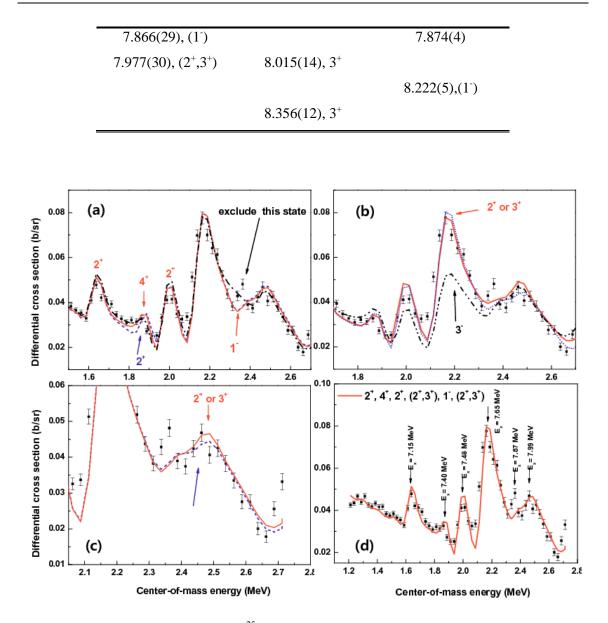


Figure 2: Excitation function for the ²⁵Al+p elastic scattering scross section fitted with an Rmatrix calculation. (a) For the second resonance around 1.89 MeV, fitting results are compared between $J^{\pi} = 2^+$ and 4^+ , and the fit for $J^{\pi} = 4^+$ gives a better result. A small bump around 2.35 MeV having an assignment of $J^{\pi} = 1^-$ resulted in an improved fit of the tail of the fourth peak. No other assignments can reproduce the tail. (b) The fourth resonance at 2.14 MeV was fitted with $J^{\pi} = 2^+$, 3^+ or 3^- assignments. Both $J^{\pi} = 2^+$ and 3^+ were consistent with the experimental data, but for $J^{\pi} = 3^-$, it totally deviated from the data. (c) High energy part of the excitation function with calculations assuming $J^{\pi} = 2^+$ or 3^+ for the last resonant peak shows that there is no difference between two assignments. (d) Final result as the best fit for $J^{\pi} = 2^+$, 4^+ , 2^+ , $(2^+$, 3^+), (1^-) , and $(2^+$, $3^+)$ are shown but without firm spin-parity assignments for the fourth and fifth resonant peaks.

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References

- [1] R. Diehl et al., Astron. And Astrophys. 298 (1995) 445.
- [2] R. A. Ward and W. A. Fowler, Astrophys. J. 238 (1980) 266.
- [3] A. Coc, M.-G. Porquet, and F. Nowacki, Phys. Rev. C 61 (1999) 015801.
- [4] R. C. Runkle, A. E. Champagne, and J. Engel, Astrophys. J. 556 (2001) 970.
- [5] A. Matic et al., Phys. Rev. C 82 (2010) 025807.
- [6] A. E. Champagne and M. Wiescher, Annu. Rev. Nucl. Part. Sci. 42 (1992) 39.
- [7] J. L. Fisker, H. Schatz, and F.-K. Thielemanm, Astrophys. J. Suppl. Ser. 174 (2008) 261.
- [8] S. Kubono et al., Eur. Phys. J. A 13 (2002) 217.
- [9] Y. Yanagisawa et al., Nucl. Instrum. Methods Phys. Res. Sect. A 539 (2005) 74.
- [10] H. S. Jung et al., Phys. Rev. C 85, (2012) 045802.
- [11] W. Benenson et al., Phys. Rev. C 15 (1977) 1187.
- [12] N. M. Larson, ORNL/TM-9179/R5, 2000 (unpublished).
- [13] A. Matic et al., Phys. Rev. C 82, (2010) 025807.
- [14] D. W. Bardayan et al., Phys. Rev. C 65, (2002) 032801.
- [15] Y. Parpottas et al., Phys. Rev. C 70, (2004) 065805.
- [16] J. Chen et al., Phys. Rev. C 85, (2012) 015805.